

A

nalisi dati con Python



R elatori

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Perché Python?

- Open Source
- Librerie ottimizzate
- Analisi dati più efficace (grande mole e grafici più precisi)



```
Python 3.6.5 (default, Apr  1 2018, 05:46:30)
[GCC 7.3.0] on linux
Type "help", "copyright", "credits" or "license" for more information.
>>> import numpy as np
>>> data=np.genfromtxt('TORI-04from2018-04-05to2018-04-06.csv',delimiter=',')
>>> data
array([[ 1.00000000e+00,  3.55270176e+08,  1.82829530e+07, ...,
         1.05321770e+02, -1.00000000e-01,  9.82000000e+02],
       [ 1.00000000e+00,  3.55270176e+08,  2.55473880e+07, ...,
         1.00918465e+02,  7.26400000e-03,  9.82000000e+02],
       [ 1.00000000e+00,  3.55270176e+08,  2.84015920e+07, ...,
         1.21021515e+02,  2.85400000e-03,  9.82000000e+02],
       ...,
       [ 2.70000000e+01,  3.55293331e+08,  1.54089992e+08, ...,
         1.07391197e+02,  3.03000000e-03,  9.82000000e+02],
       [ 2.70000000e+01,  3.55293331e+08,  2.45941266e+08, ...,
         1.01238701e+02,  9.18510000e-02,  9.82000000e+02],
       [ 2.70000000e+01,  3.55293331e+08,  2.52633024e+08, ...,
         1.04087105e+02,  6.69200000e-03,  9.82000000e+02]])
>>> 
```

Librerie

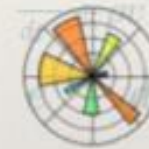
- ❖ Installazione delle librerie da prompt comandi: `pip install nomelibreria` (vedi slide seguente.)



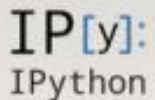
NumPy
Base N-dimensional array package



SciPy library
Fundamental library for scientific computing



Matplotlib
Comprehensive 2D Plotting



IPython
Enhanced Interactive Console



Sympy
Symbolic mathematics



pandas
Data structures & analysis

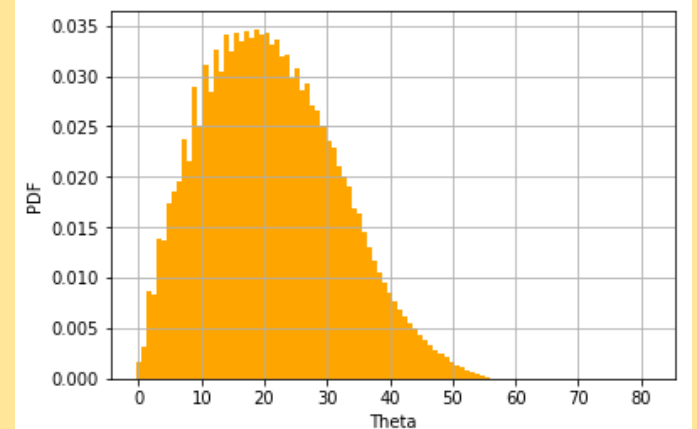
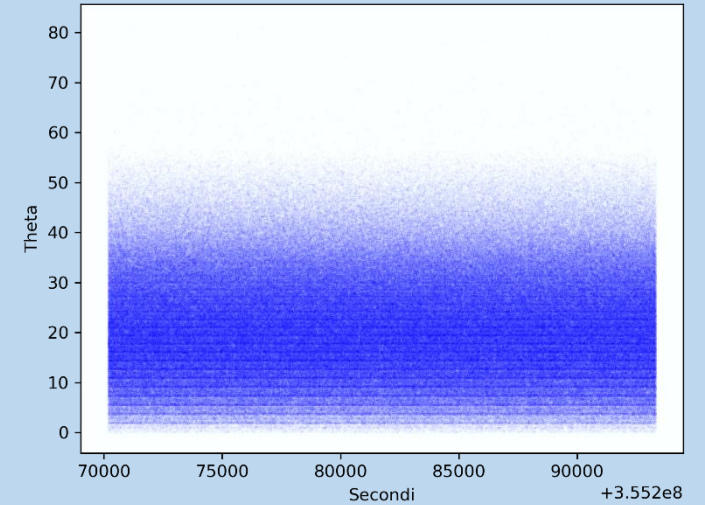
- ❖ Le librerie contengono codice compilato scritto in c++. Alcune librerie richiedono il compilatore installato sulla macchina al fine di ottimizzare le prestazioni di calcolo.

Distribuzione angolare

Angolo Theta

```
#Scatter
import numpy as np
import matplotlib.pyplot as plt
data=np.genfromtxt("TORI-04from2018-04-05to2018-04-06.csv",delimiter=",")
x=(data[1:,1]+data[1:,2]*10**(-9)) #secondi + nanosecondi
y=data[1:,3] #Theta
plt.scatter(x,y,0.0005)
plt.xlabel( "Secondi" )
plt.ylabel( "Theta" )
plt.show()

#Istogramma
Nc=100 #numero di intervalli
pdf, bins = np.histogram( y, Nc, density = True )
nextr = len( bins )
dx = bins[1] - bins[0]
plt.xlabel( "Theta" )
plt.ylabel( "PDF" )
plt.grid()
plt.bar( bins[:nextr-1], pdf, dx, color = "orange" )
plt.show()
```



Tempo di volo

```
import scipy.optimize as syt
import numpy as np
def F( x, mu, sigma ):
    ff = 1/(np.sqrt( 2.0 * np.pi ) * sigma ) * np.exp(-(x-mu)**2.0)/(2.0*sigma**2))
    return ff
imin = 0
imax = 999
Nc=1000
pdf, bins = np.histogram( y, Nc, density = True )
nexttr = len( bins )
xc = 0.5 * ( bins[0:nexttr-1] + bins[1:nexttr] )
xdata = xc[imin:imax]
ydata = pdf[imin:imax]
par, corr = syt.curve_fit( F, xdata, ydata )
nexttr = len( bins )
dx = bins[1] - bins[0]
plt.xlabel( r"$x$" )
plt.ylabel( "Pdf" )
plt.bar( bins[:nexttr-1], pdf, dx, color = "g" )
yc = ( F( xc, par[0], par[1]) )
plt.plot( xc, yc, "r-o" )
plt.xlim(0,8)
plt.title( '$\mu$='+str(par[0])+' , $\sigma$='+str(par[1]), fontsize=10)
plt.show()
```

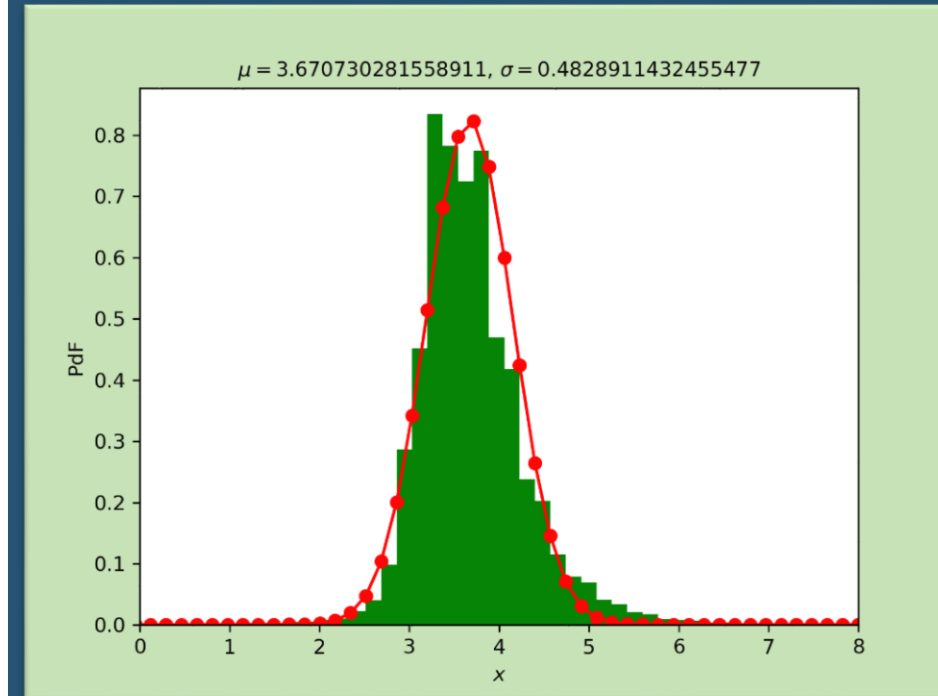


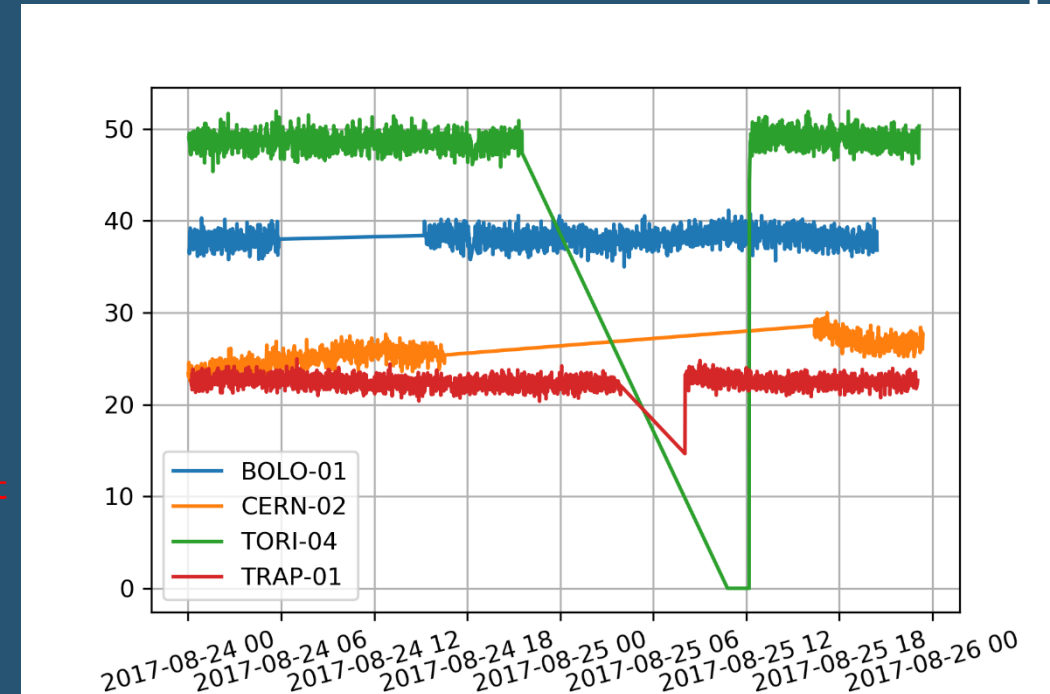
Grafico dei dati con curva di fit gaussiano

Confronti tra osservatori

```
import numpy as np
import matplotlib.pyplot as plt
import matplotlib.dates as md
import datetime as dt

data_b1=np.genfromtxt("BOLO-01.csv",delimiter=",")
data_c2=np.genfromtxt("CERN-02.csv",delimiter=",")
data_t4=np.genfromtxt("TORI-04.csv",delimiter=",")
data_t1=np.genfromtxt("TRAP-01.csv",delimiter=",")
t0 = 1167609600 #la differenza di secondi tra 1/1/1970 e 1/1/2007
c = 0
data_b1= data_b1[data_b1[:,c].argsort()]
nxi_b1=data_b1.shape[0]
data_b1 = np.unique(data_b1,axis=0)
nx_b1=data_b1.shape[0]
timestamps_b1 = data_b1[:,0] + t0
xts_b1 = md.date2num([dt.datetime.fromtimestamp(s) for s in timestamps_b1])
y_b1 = data_b1[:,4] # estraggo la colonna dei conteggi/minuto RateTrackEvent
[...]
plt.xticks( rotation=15 )
ax=plt.gca()
#xfmt = md.DateFormatter('%Y-%m-%d %H:%M:%S')
xfmt = md.DateFormatter('%Y-%m-%d %H')
ax.xaxis.set_major_formatter(xfmt)
p_b1, =plt.plot(xts_b1,y_b1)
p_c2, =plt.plot(xts_c2,y_c2
```

```
p_t4, =plt.plot(xts_t4,y_t4)
p_t1, =plt.plot(xts_t1,y_t1)
plt.legend([p_b1, p_c2, p_t4,p_t1],["BOLO-01","CERN-02","TORI-04","TRAP-01"])
plt.grid()
#plt.savefig('Multiple.png',dpi=300)
plt.show()
```



Download multiplo

```
import urllib.request
#gmin: primo giorno da scaricare
#gmax: ultimo giorno da scaricare
#tel: nome strumento
#Es. download(1,30,"CERN-02")

def download(gmin,gmax,tel):
    f=open("lista", "a+")
    for i in range(gmin+1,gmax+1):
        if i<10:
            a1="0"+str(i-1)
            a2="0"+str(i)
        elif i==10:
            a1="0"+str(i-1)
            a2=str(i)
        else:
            a1=str(i-1)
            a2=str(i)
        name=tel+'_2017-08-'+a1+'_2017-08-'+a2+'_summary_Trending.csv'
        url = 'https://iatw.cnaf.infn.it/eee/monitor/dqmreport2/'+tel+'/'+'2017-08-'+a2+'/'+'name
        urllib.request.urlretrieve(url,name)
        print("Scaricato:"+name)
        f.write(name+ '\n')
    f.close()
```

Cartella contenente 60 files giornalieri

21:49	443.873	CERN-02_2017-08-17_2017-08-18_summary_Trending.csv
21:49	449.249	CERN-02_2017-08-18_2017-08-19_summary_Trending.csv
21:49	449.753	CERN-02_2017-08-19_2017-08-20_summary_Trending.csv
21:49	439.505	CERN-02_2017-08-20_2017-08-21_summary_Trending.csv
21:49	442.361	CERN-02_2017-08-21_2017-08-22_summary_Trending.csv
21:49	436.985	CERN-02_2017-08-22_2017-08-23_summary_Trending.csv
21:49	380.537	CERN-02_2017-08-23_2017-08-24_summary_Trending.csv
21:49	224.633	CERN-02_2017-08-24_2017-08-25_summary_Trending.csv
21:49	286.457	CERN-02_2017-08-25_2017-08-26_summary_Trending.csv
21:49	444.377	CERN-02_2017-08-26_2017-08-27_summary_Trending.csv
21:49	447.905	CERN-02_2017-08-27_2017-08-28_summary_Trending.csv
21:49	448.577	CERN-02_2017-08-28_2017-08-29_summary_Trending.csv
20:46	<DIR>	Codice
22:14	320.393	TORI-04_2017-08-01_2017-08-02_summary_Trending.csv
22:14	403.217	TORI-04_2017-08-02_2017-08-03_summary_Trending.csv
22:14	439.001	TORI-04_2017-08-03_2017-08-04_summary_Trending.csv
22:14	452.105	TORI-04_2017-08-04_2017-08-05_summary_Trending.csv
22:14	451.937	TORI-04_2017-08-05_2017-08-06_summary_Trending.csv
22:14	433.289	TORI-04_2017-08-06_2017-08-07_summary_Trending.csv
22:14	423.041	TORI-04_2017-08-07_2017-08-08_summary_Trending.csv
22:14	439.841	TORI-04_2017-08-08_2017-08-09_summary_Trending.csv

Unire più file

```
import numpy as np
f=open('lista','r')
flst=f.readlines()
print('Nella lista ci sono '+str(len(flst))+ ' file')
if (len(flst)>1):
    print('Importo il file ',1,' ->',flst[0][0:-1])
    data=np.genfromtxt(flst[0][0:-1],delimiter=',')
    print('Le dimensioni della matrice dati:',data.shape)
    for i in np.arange(1,len(flst)-1,1):
        print('Importo il file ',i+1,' ->',flst[i][0:-1])
        data=np.vstack((data,np.genfromtxt(flst[i][0:-1],delimiter=',')))
        print('Le dimensioni della matrice dati è diventata:',data.shape)
else:
    print('Importo unico file ->',flst[0:-1])
    data=np.genfromtxt(flst[0:-1],delimiter=',')
    print('Le dimensioni della matrice dati:',data.shape)
print('Sto ordinando la matrice dati in funzione della prima colonna...')
c = 0
data= data[data[:,c].argsort()]
print('...e adesso elimino le righe duplicate...')
nxi=data.shape[0]
data = np.unique(data,axis=0)
nx=data.shape[0]
print('Ho eliminato complessivamente ',nxi-nx,' righe')
print('La nuova matrice dati ha le seguenti dimensioni',data.shape)
```

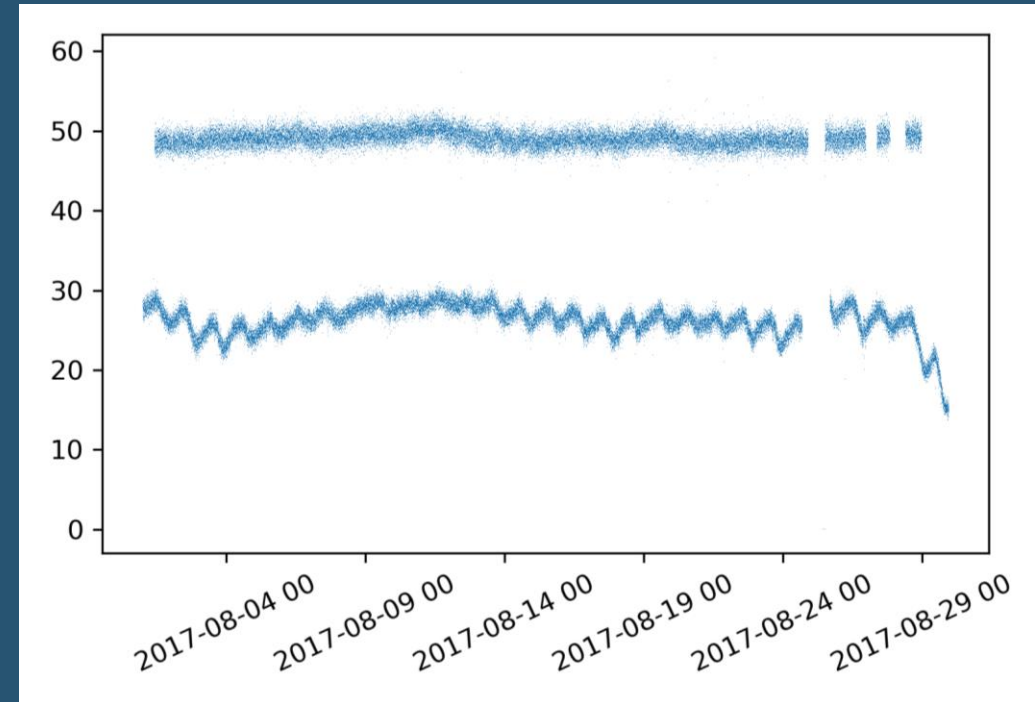
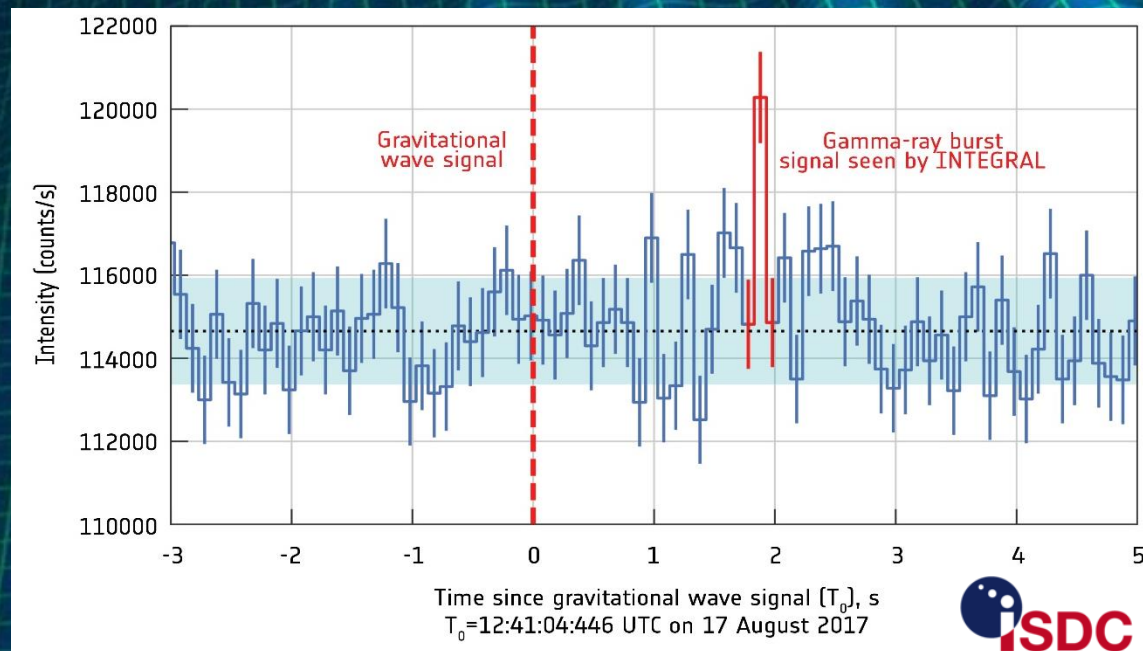


Grafico mensile su TORI-04 e CERN-02 (RateTrackEvent)

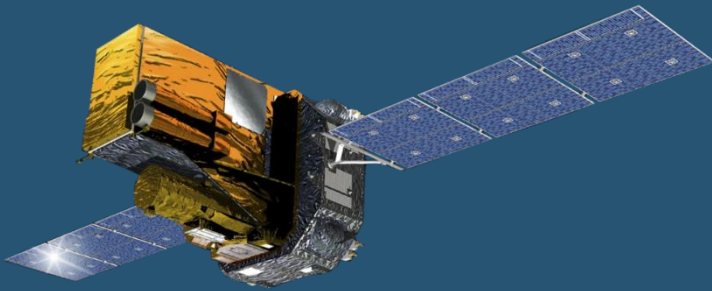
Correlazione tra flusso di raggi cosmici e fusione di due stelle a neutroni

17 agosto 2017

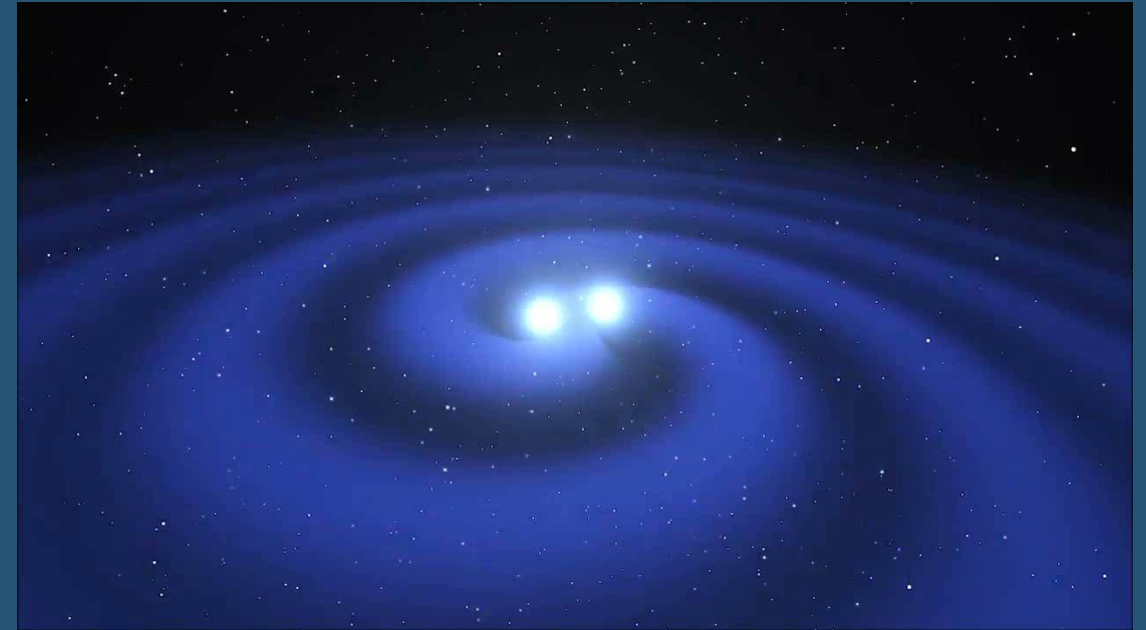
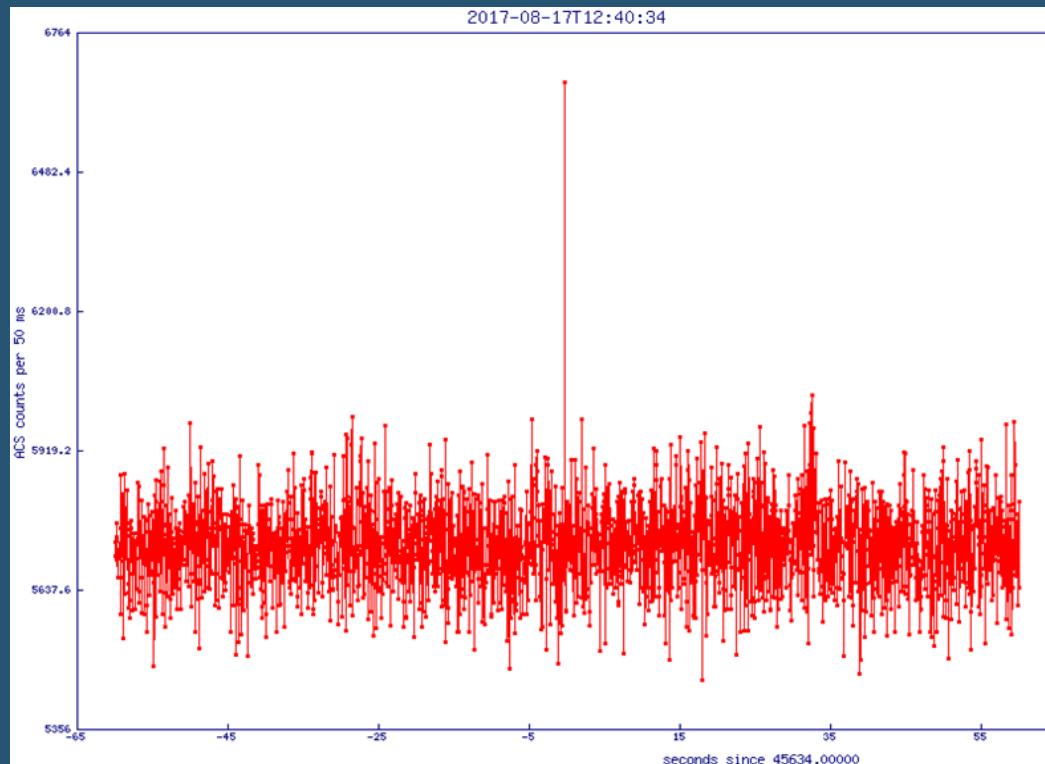


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Prof. Vincenzo Di Matteo

INTEGRAL: INTErnational Gamma-Ray Astrophysics Laboratory



Lampo gamma dopo le onde gravitazionali



```
import numpy as np
f=open('dataintegral.txt','r')
data_integral=np.genfromtxt(f,delimiter=" ")
```

```
import matplotlib.pyplot as plt
x=data_integral[:,0]
y=data_integral[:,1]
plt.plot(x,y)
plt.show()
```

Δt tra fotoni e protoni emessi simultaneamente dalla kilonova

$$m = 10^{20} \text{ eV} \quad d = 1,3 \cdot 10^8 \text{ a.l.}$$
$$m_0 = 10^9 \text{ eV}$$

$$\gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} = 10^{11} \quad \longrightarrow \quad v = c(1 - \gamma^{-2})^{\frac{1}{2}}$$

$$\Delta t = \frac{d}{v} - \frac{d}{c} = \frac{d}{c(1 - \gamma^{-2})^{\frac{1}{2}}} - \frac{d}{c} = \frac{d}{c} \left[(1 - \gamma^{-2})^{-\frac{1}{2}} - 1 \right] \approx \frac{d}{c} \left[1 - \left(-\frac{1}{2} \right) \gamma^{-2} - 1 \right] = \frac{d}{c} \frac{1}{2} \gamma^{-2}$$

$$\frac{d}{c} = 1,3 \cdot 10^6 \text{ anni} = (1,3 \cdot 10^6) \cdot (3,1 \cdot 10^7 \text{ s}) = 4,03 \cdot 10^{15} \text{ s}$$

$$\Delta t = \left(\frac{4,03 \cdot 10^{15}}{2} \text{ s} \right) \gamma^{-2} = (2,02 \cdot 10^{15} \text{ s}) \gamma^{-2}$$

γ^2	$\Delta t \text{ (s)}$
10^{11}	$2,02 \cdot 10^7$
10^8	0,202
10^7	20,2
10^6	$2,02 \cdot 10^3$
10^5	$2,02 \cdot 10^5$

$$m_1 \in (1.36 - 2.26)M_{\odot} \rightarrow \overline{m}_1 = \frac{1.36 + 2.26}{2} = 1.81M_{\odot}$$

$$m_2 \in (0.86 - 1.36)M_{\odot} \rightarrow \overline{m}_2 = \frac{0.86 + 1.36}{2} = 1.11M_{\odot}$$

$$m_{tot} = 2.82^{+0.47}_{-0.09}M_{\odot}$$

$$\overline{m}_1 + \overline{m}_2 = 2.92M_{\odot}$$

$$E = (\overline{m}_1 + \overline{m}_2) - m_{tot} = 2.92 - 2.82 = 0.1M_{\odot}$$

$$1M_{\odot} = 1.989 \times 10^{30} \text{ kg} = 1.116 \times 10^{66} \text{ eV}$$

$$E(\text{eV}) = 1.116 \times 10^{66} \text{ eV} \cdot 0.1M_{\odot} \\ = 1.116 \times 10^{65} \text{ eV} \quad \text{Energia emessa nella coalescenza}$$

$$n_{p^+} = \frac{E}{m_{p^+}} = \frac{1.116 \times 10^{65} \text{ eV}}{9.421 \times 10^8 \text{ eV}} = 0.118 \times 10^{57} \quad \text{Numero di protoni emessi}$$

$$d = 130 \text{ mln a.l.} = 1.3 \times 10^8 \text{ a.l.}$$

$$1 \text{ a.l.} = 9.461 \times 10^{15} \text{ m}$$

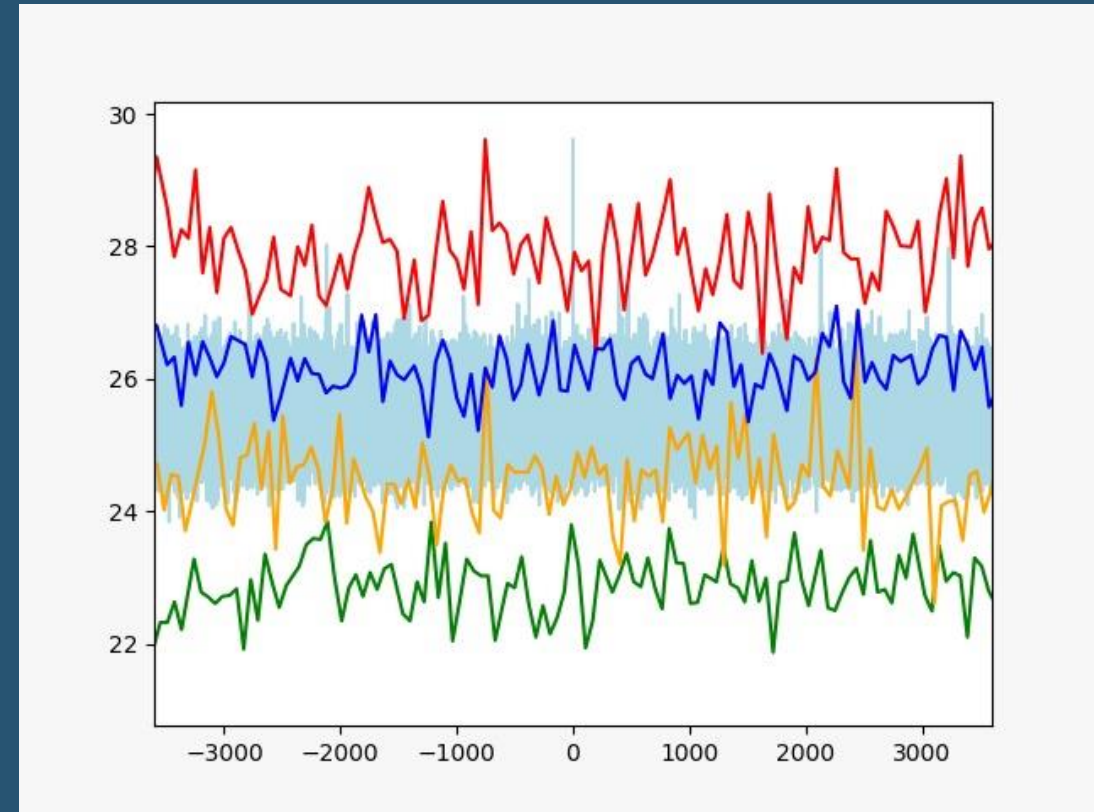
$$d(\text{m}) = 9.461 \times 10^{15} \cdot 1.3 \times 10^8 = 12.229 \times 10^{23} = 1.229 \times 10^{24} \text{ m}$$

$$\frac{n_{p^+}}{m^2} = \frac{n_{p^+}}{4\pi d^2} = \frac{0.118 \times 10^{57}}{4\pi(1.229 \times 10^{24} \text{ m})^2} = \frac{0.118 \times 10^{57}}{1.899 \times 10^{49} \text{ m}^2} \\ = 6.2 \times 10^6 \frac{n_{p^+}}{m^2} \quad \text{Numero di protoni per m}^2 \text{ sulla terra}$$

<http://iopscience.iop.org/article/10.3847/2041-8213/aa91c9/meta>

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Confronto tra osservatori:

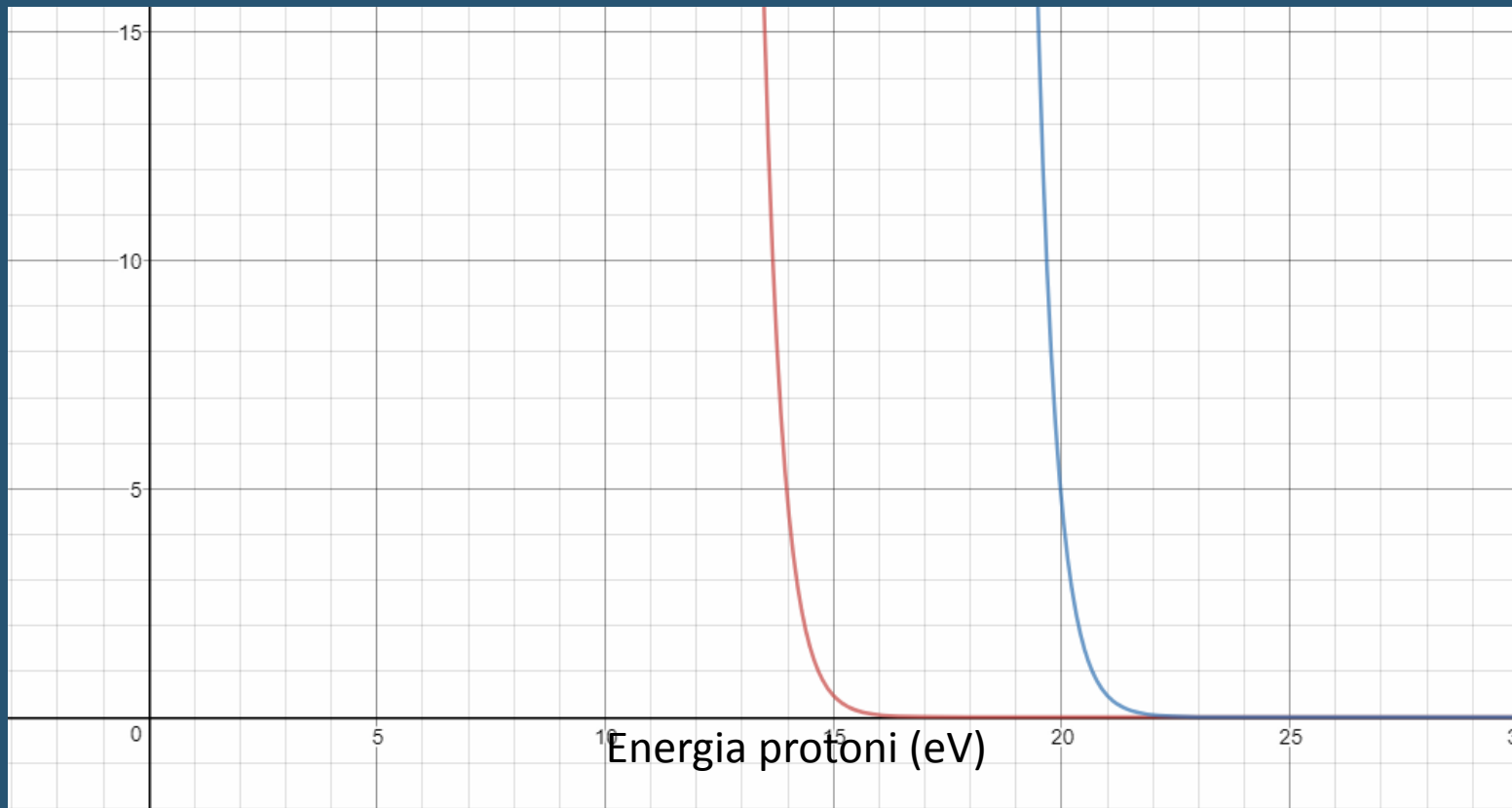
-Background **INTEGRAL**;

-**Rosso** **CERN-02**;

-**Arancione** **SALE-02**;

-**Verde** **TORI-04**;

-**Blu** sovrapposizione tre osservatori EEE



$$y = \frac{(1.116 \cdot 10^{65})}{(10^x) \cdot (4\pi) \cdot (1.899 \cdot 10^{49})}$$



$$y = \frac{(1.116 \cdot 10^{65})}{(10^x) \cdot (4\pi) \cdot (1.899 \cdot 10^{43})}$$

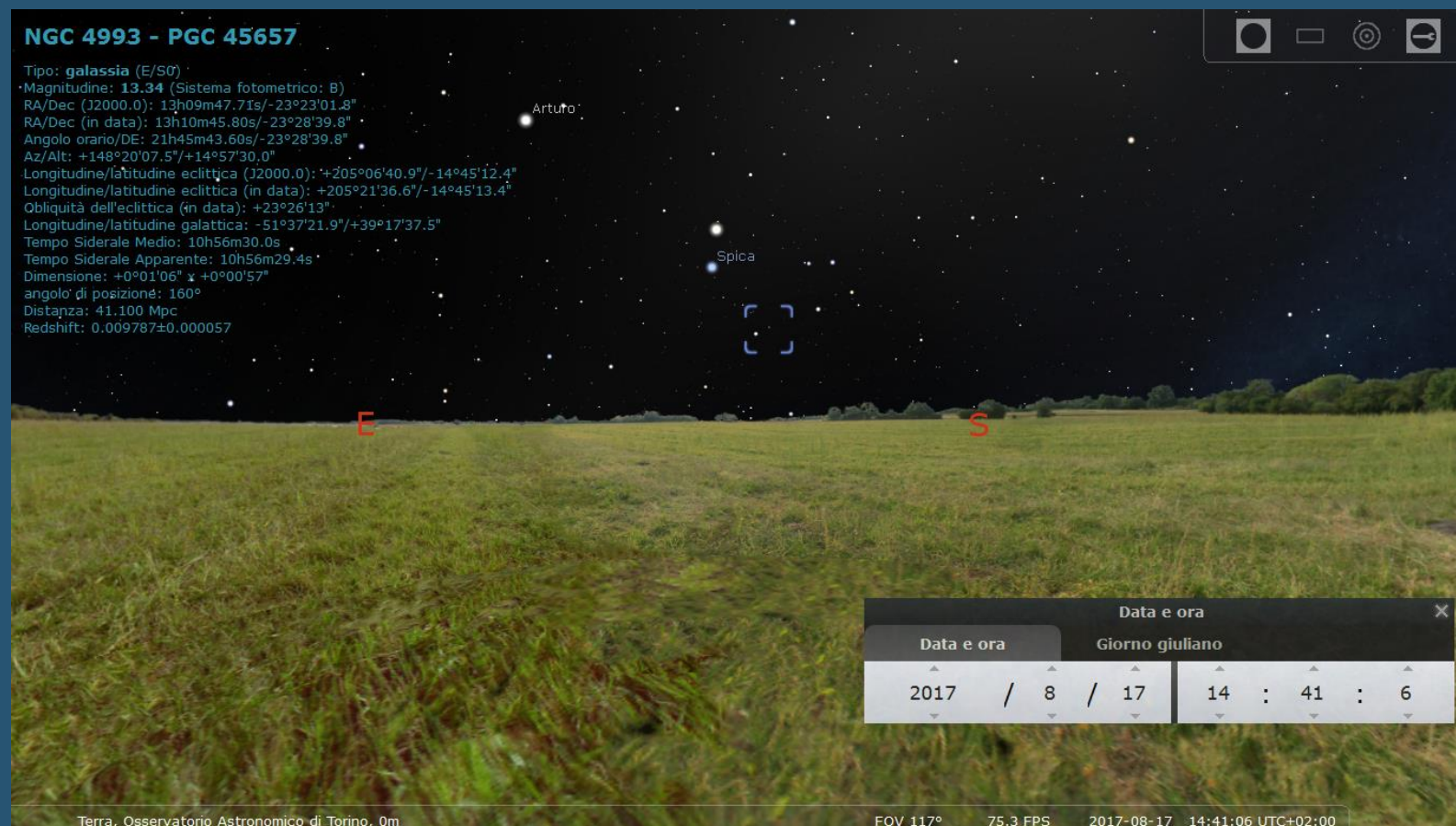
$$y = \frac{E_{emessa}}{E_{p+} 4\pi d^2}$$

Numero di protoni con $E = 10^{20}$ eV : **4.7×10^{-6} per m^2**
 4.5 per km^2

Attivi solo 7 telescopi su 52 !!

Di cui sei ad alte latitudini

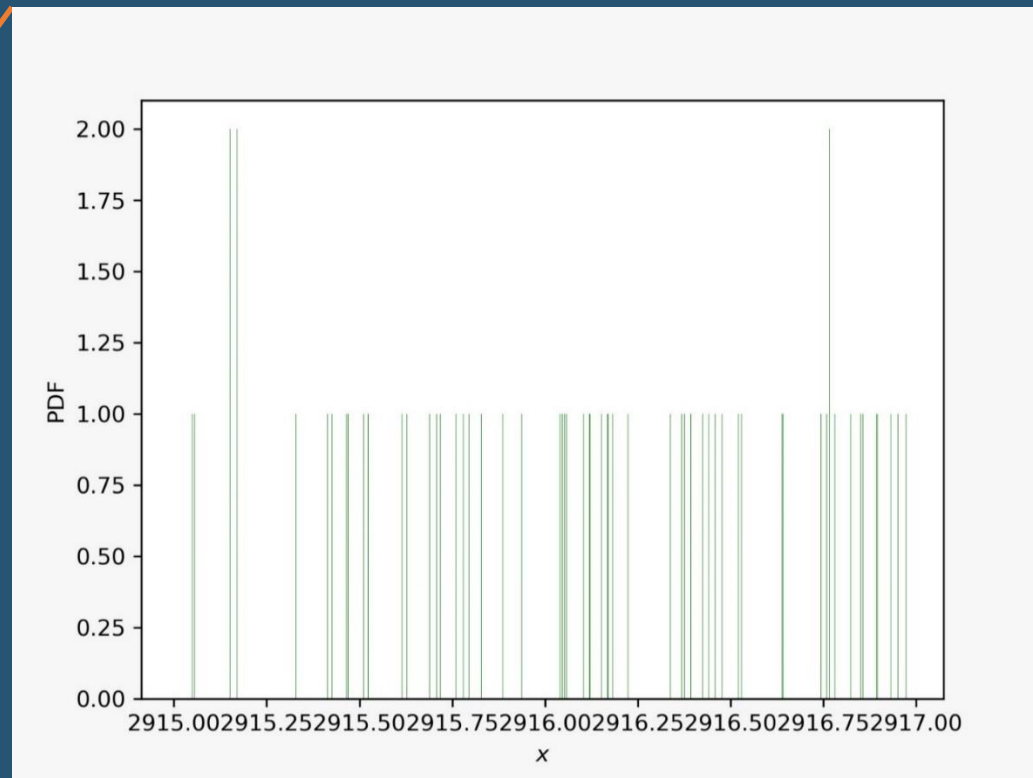
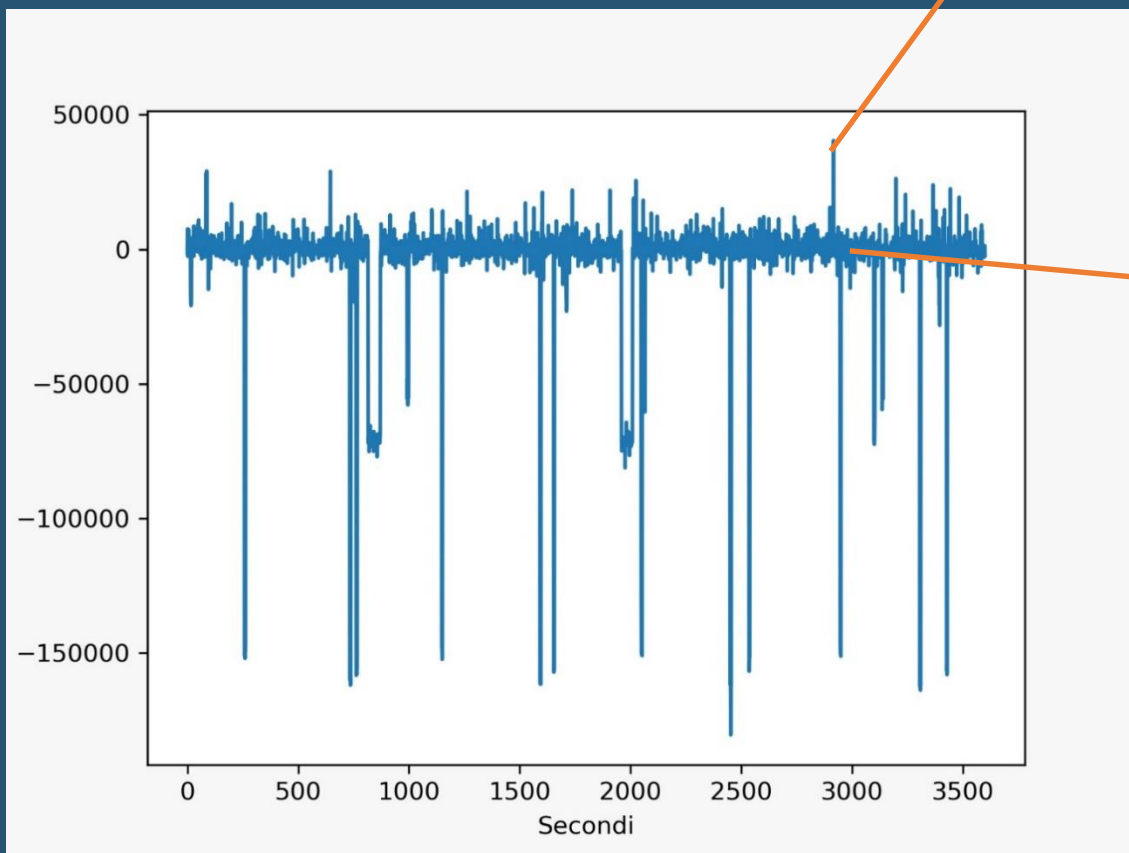
Rilevatore	Altezza della NGC 4993
AREZ-01	17°50'37"
BOLO-04	16°45'17"
CERN-01	13°22'30"
CERN-02	13°22'30"
SALE-02	21°17'43"
TORI-03	14°57'43"
TORI-04	14°57'43"



Osservatorio di Pino Torinese

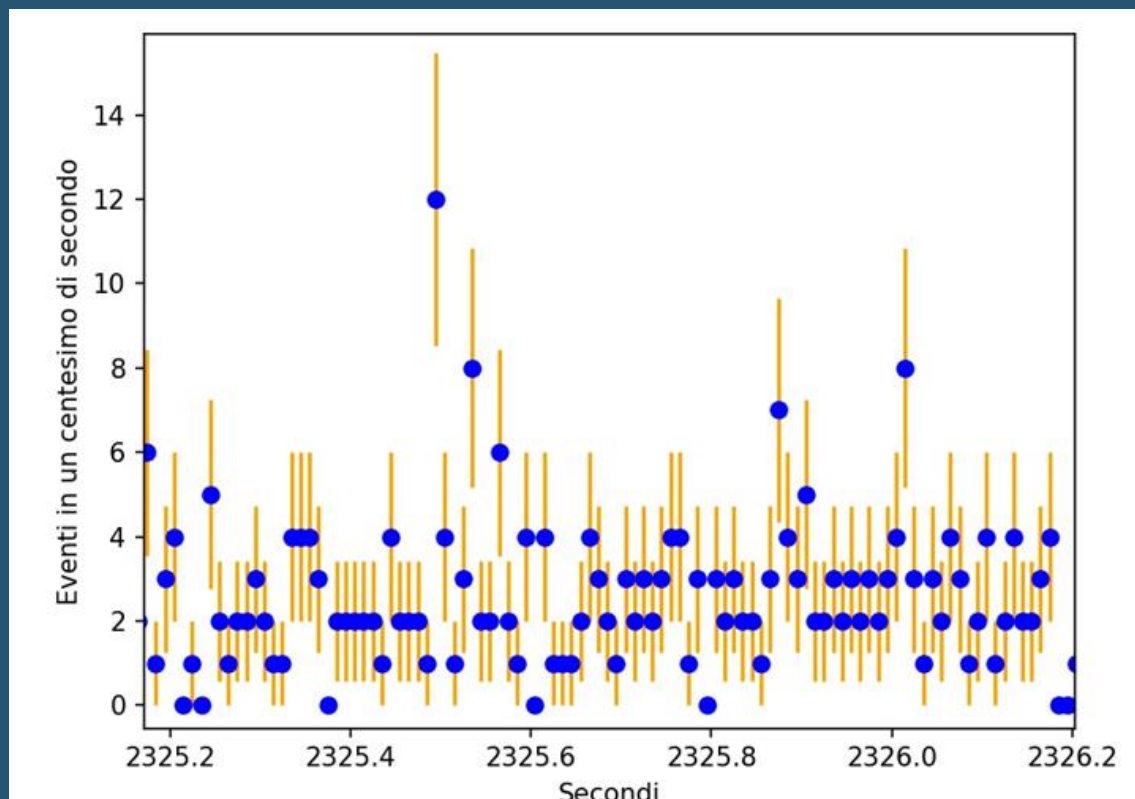
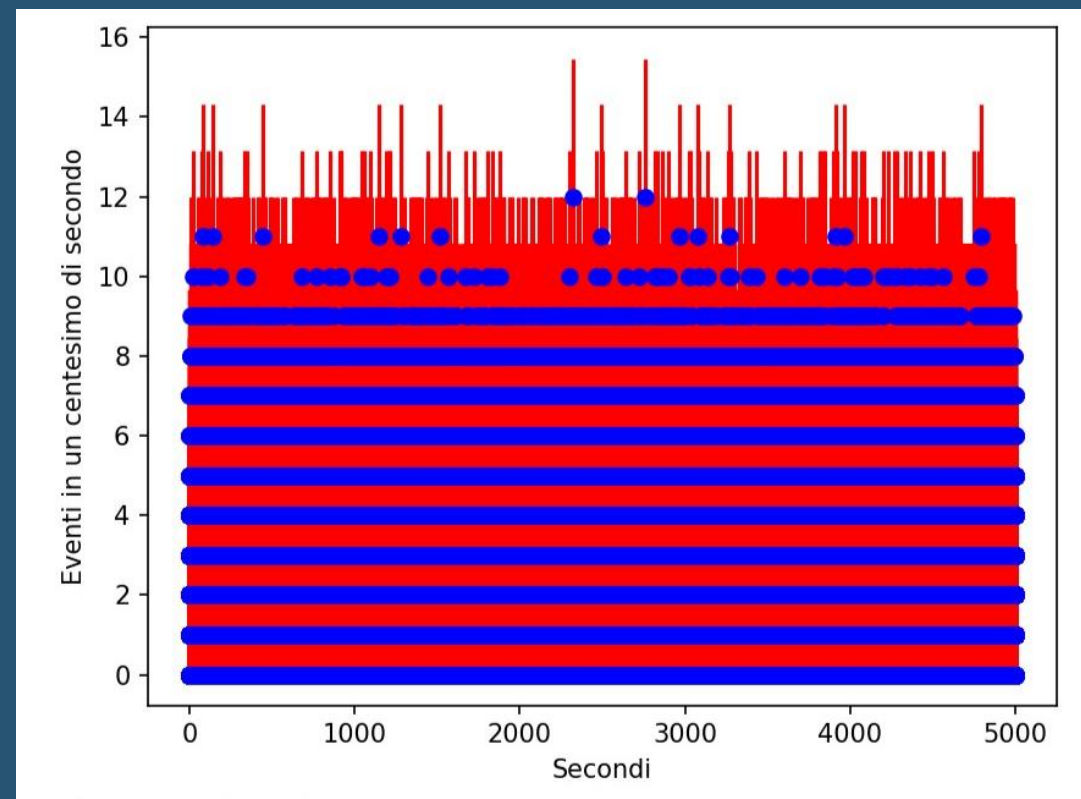
Somma di tutti gli eventi (passaggi di muoni)
rilevati da sei dei sette osservatori (si esclude
Salerno per problemi di importazione dati)

$$\sqrt[3]{\sum_i (y_n - \bar{y}_i)^3}$$



Eventi al microsecondo interni al secondo del picco più alto

Somma degli eventi detetti dai sei rivelatori per ogni centesimo di secondo su un range di 5000 s



Somma degli eventi detetti dai sei rivelatori per ogni centesimo di secondo su un range di 1 s

Lavori in corso:

- **Determinazione della velocità del vento solare in relazione al tipo di brillamento solare. (X9 e X2)**
- **Determinazione della variazione del rapporto del flusso di muoni nel tempo dovuto all'effetto giorno-notte**
- **Confronto tra conteggi/tempo nei casi di Sole attivo e Sole quieto attraverso due diversi telescopi: EEE (muoni) e OULU (neutroni)**

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- **Dott. Ivan Gnesi** – INFN Torino
- **Prof. Giovanni Peres** – Dipartimento di Fisica Palermo (UNIPA)
- **Dott. Carlo Ferrigno** – INTEGRAL Science Data Centre (University of Geneva)
- **Prof. Antonio Maggio** – Osservatorio Astronomico di Palermo (OAPA)
- **Dott.ssa Paola La Rocca** – INFN Catania

"Remember to look up at the stars and not down at your feet. Try to make sense of what you see and wonder about what makes the universe exist. **Be curious.** And however difficult life may seem, there is always something you can do and succeed at. It matters that you don't just give up."

-Stephen Hawking

