

Updates on POLA-R rate time series analysis

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Rate POLA-R a Ny Ålesund

- Alcune delle analisi in corso:
 - Analysis of the may 2024 Forbush decrease at Ny Ålesund, with BSS, CRNS and POLA (and Barentsburg)
 - Submitted to 'Advances in Space Research'
 - Invited presentation at EGU2025 (next slide)
 - to be submitted to ICRC (who?)
 - correzioni per temperatura
 - analisi di periodicità

EGU2025, Vienna 27 aprile – 2 maggio

- Observation of the Forbush decrease during the May 2024 solar storms with different muon and neutron detectors in the high-latitude site of the Svalbard archipelago
by Ombretta Pinazza et al.

accepted for session G14.4 – Cosmic rays across scales and disciplines: the new frontier in environmental

Nuova proposta di collaborazione a Ny Ålesund

- Sandipan Dawn, from Bhabha institute, Mumbai (India)
- Paper: Characterizing interplanetary magnetic field fluctuations at arctic using cosmic ray secondaries—An approach with machine learning
 - DOI: [10.1016/j.astropartphys.2025.103087](https://doi.org/10.1016/j.astropartphys.2025.103087)
 - *Astropart.Phys.* 167 (2025) 103087
- Plans: deploy a combination of plastic scintillator-based muon detectors along with He-3-based neutron detectors

POLA-R rate : cuts and corrections

- quality cuts: status flag, missing pressure

diff(rate,ratej) <= 3 Hz

tempELE max = 40.0 °C

rate_min = 10.0 Hz

pseff_min = 0.2

- pseudoefficiency correction

rate = sum(ratePair_j/effAv)

effAv = $e_1 * e_2 * e_3 + e_1 * e_2 * e_4 + e_1 * e_3 * e_4 + e_2 * e_3 * e_4 - 3 * e_1 * e_2 * e_3 * e_4$

- pressure correction

rateC = rate / math.exp(beta*(pres-presAvg))

beta <= curve_fit(Linear, P-<P>, 100*ln(R/<R>))

beta ~ -0.00236 evaluated on the whole period

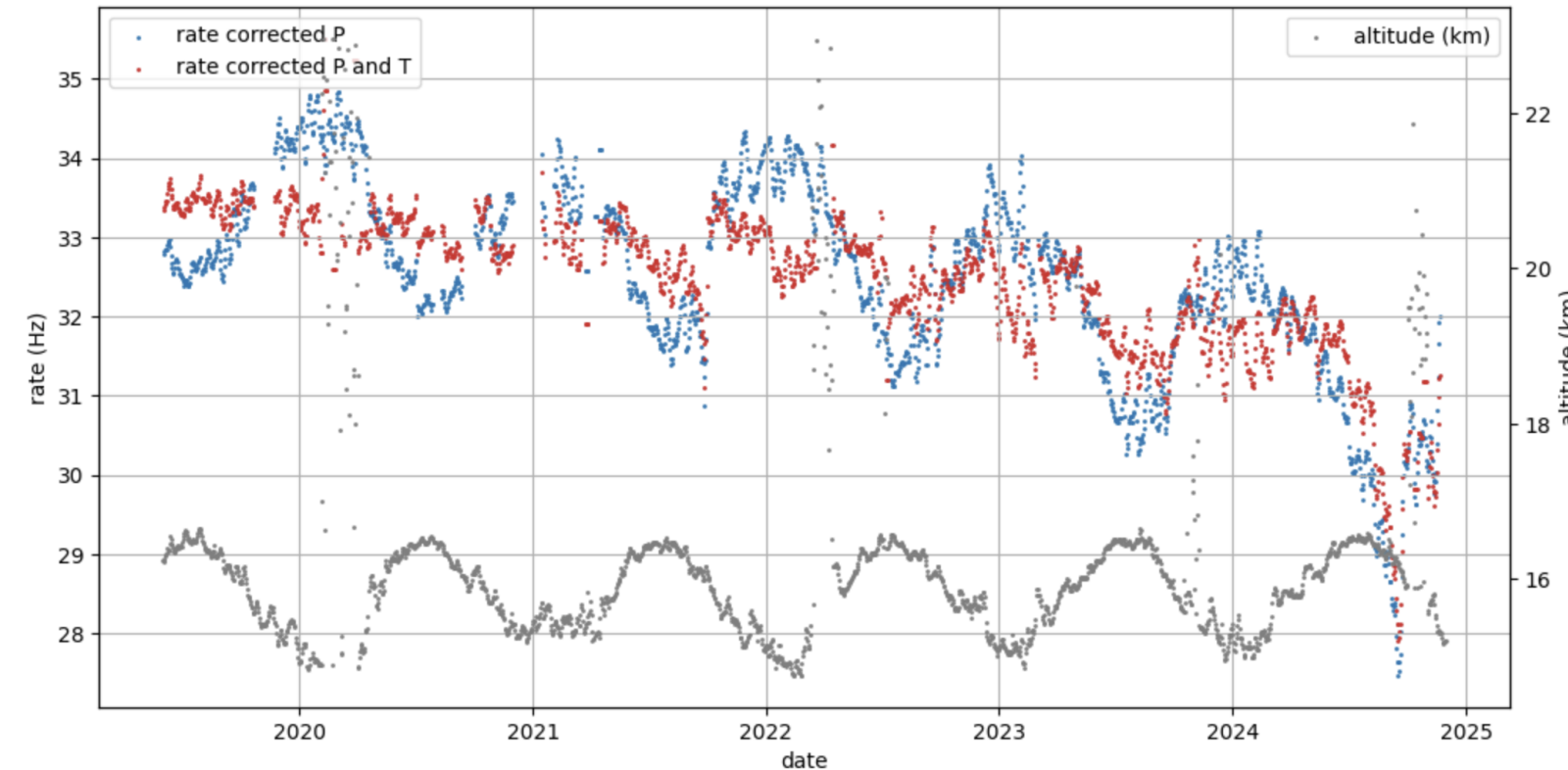
$$\ln \left[\frac{I_{[x,y]}(t)}{\langle I_{[x,y]} \rangle} \right] * 100\% = \beta_{[x,y]} * [P(t) - \langle P \rangle],$$

POLA-R rate : temperature correction studies

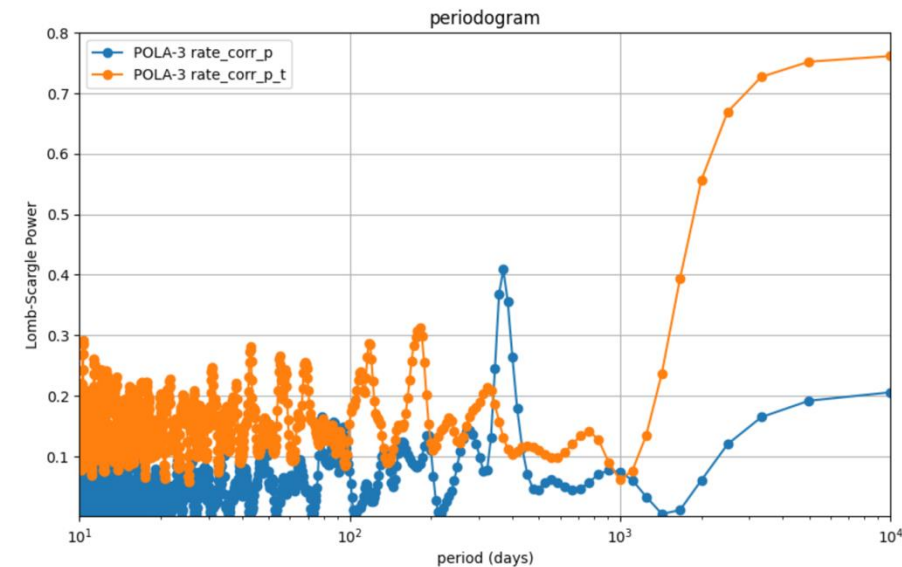
- de Mendonça papers 2016-2019 : 7 methods

- 1. atmospheric expansion method $\Delta I_T = \alpha_{ATE} * \Delta H[p], p = 100 \text{ hPa}$

POLA-3 rate corrected for pressure and temperature - method1



Pearson before = -0.634
Pearson after = -0.084



Appendix

‘Methods for describing and removing the temperature effect’

Rafael R. S. de Mendonça et al.

- 7 methods:
 1. Atmospheric expansion method (altitude profile at $p=100$ hPa)
 2. Ground method (based on the temperature near the ground)
 3. **MMP method (based on temperature variation at the MMP altitude level)**
 4. Combination of 1+2+3
 5. THM: Theoretical model (integral of temperature effects at different levels)
 6. **MSS: Mass Weighted method (approx. of THM, as the sum of weighted contributions at different atm levels)**
 7. Effective temperature (single parameter related to pion and nucleons absorption in the atm., no annual periodicity)

MMP = altitude at which we see maximum production of secondary CR, estimated at 16.5 km, mainly pion decay

BLUE = negative correlation

RED = positive correlation

GREY = no periodicity

BOLD = preferred

Correlation rateCP and rateCPT with envTemp

Correlation between the corrected rate and external temperature at 2m (CCT)

