# Why precision timing for <br> Extreme Energy Events 

Giornate di Studio di Extreme Energy Events

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## The first extensive shower event candidate at the EEE observatory



The time displacement among the 3 telescopes was
~ 1 microsecond
Cuts on the relative angles among tracks were applied

EEE has both
timing + tracking
capabilities!

## An extensive air shower candidate: how to detect it? Time plays a fundamental role

To "detect" a candidate EEE has to define a coincidence time window among telescopes.

The window has to be as large as to include all possible shower direction

Thus, in principle the minimum coincidence window is

L/v
and if
$\mathbf{L \sim 1} \mathbf{~ k m}$ and $\mathbf{v \sim c}$

$\Delta T \sim 3$ microseconds!

## Can we correct some of the sistematic uncertainties?

using a "guess" on the shower direction we can try to
correct the time difference between detectors
by subctracting to the Telescope 2 particle time the additional "light path"
$\Delta L \cos (0) / v$


## Physics is always more complex than what one would think....

There are other uncertainties sources.
They are not just sistematic uncertainties

- disc thickness
- disc shape
due to the complex particle cascade process particles are not travelling in time

Disc thickness depends on energy of the primary particle

Disc shape is not linear (parabolic)


## Examples of what we get after corrections

CERN (15 m)
correction assuming $\Delta \phi=-2.35, \Delta \mathrm{~L}=15.0 \mathrm{~m}$


## LAQU (204 m)

correction assuming $\Delta \phi=-2.86, \Delta \mathrm{~L}=204.0 \mathrm{~m}$


## Examples of what we get after corrections

## CAGL (520 m)

correction assuming $\Delta \phi=1.26, \Delta \mathrm{~L}=520.0 \mathrm{~m}$


SAVO (1182 m)
correction assuming $\Delta \phi=-0.33, \Delta \mathrm{~L}=1182.0 \mathrm{~m}$


## How much the time measurement uncertainty matters

25 ns fis +3 ns err


25 ns fis +10 ns err


25 ns fis +50 ns err


# How much the <br> time measurement uncertainty matters 

Spurious coincidences (blue area) $=2$ f1 f2 $\Delta T$
$\mathrm{N} 1=\mathrm{f} 1 \Delta \mathrm{~T}$
number of events within a time window $\Delta T$ for a given telescope with a typical event frequency f1

$N 2=f 2 \Delta T$
number of events within a time window $\Delta T$ for a given telescope with a typical event frequency f1

## How much the <br> time measurement uncertainty matters

Spurious coincidences (blue area) $=2 \mathrm{f} 1 \mathrm{f} 2 \Delta \mathrm{~T}$



## A very rare event search... where time really matters!

$\mathbf{L \sim 1 0 0 0} \mathbf{~ k m}$ and $\mathbf{v ~ c}$
$\Delta T$ ~ 3 milliseconds


## A very rare event search... where time really matters!

- Physics is not known
- a wide time
displacement between events could be driven by physics
- Very Iow statistics
- Signal/Noise not possible

In case of clock misalignement
NO EVENTS!!!


## thus .... <br> let's try to deepen our knowledge about Time....

